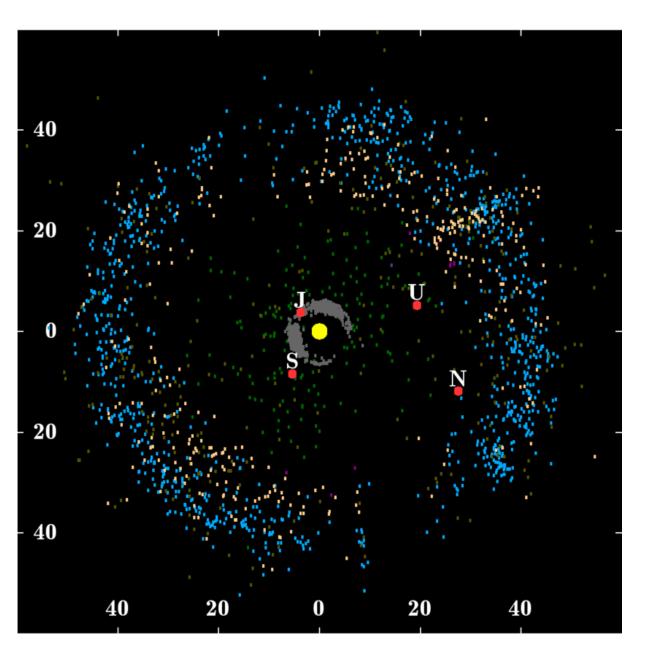
## The Solar system



### Kuiper belt (30-50 AU) KBO=Kuiper Belt Object Known objects in the Kuiper belt 티

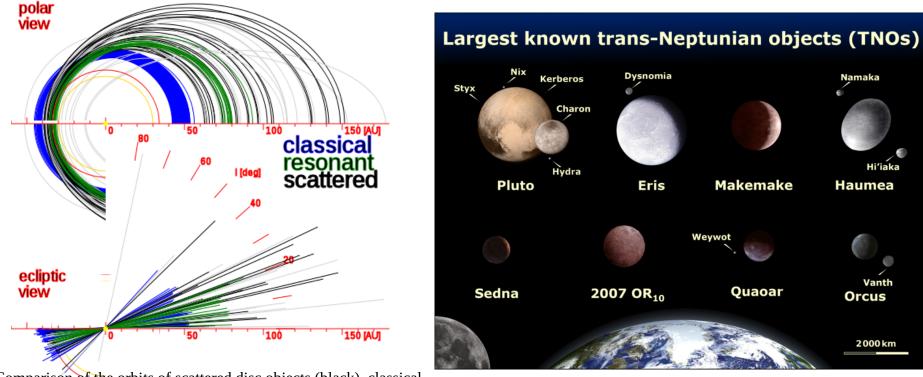


• Can be roughly divided into the **classical** belt and the **resonances**. Resonances are orbits linked to that of Neptune (e.g. twice for every three Neptune orbits, or once for every two).

• Makemake (45.79 AU average), although smaller than Pluto, is the largest known object in the **classical** Kuiper belt (KBO not in a confirmed resonance with Neptune).

• The dwarf planet Pluto (39 AU average) is the largest known object in the Kuiper belt. When discovered in 1930, it was considered to be the ninth planet; this changed in 2006 with the adoption of a formal definition of planet (3/2 **resonance** with Neptune)

# Typical objects in KBO



Comparison of the orbits of scattered disc objects (black), classical KBOs (blue), and 2:5 resonant objects (green). Orbits of other KBOs are gray. (Orbital axes have been aligned for comparison.)

### Trans-neptunian planets

- Similar to Uranus or Neptune, but formed directly at 125-250 AU?
- Formed closer and scattered by the giant planets ealry in SS history?
- Captured from another planetary system?

## Discovery of Sedna-like body Trujillo & Sheppard, Nature 507, 471 (2014)

Surveyed 52 deg<sup>2</sup> of sky for new inner Oort cloud objects using the Dark Energy Camera (DECam) at CTIO 4-m telescope: each patch imaged 3 times with 1.5 h to 2 h between images to detect the motion of objects beyond 300 AU when compared to background stars + position and colour of 2012 VP 113 measured using the Magellan 6.5-m telescope

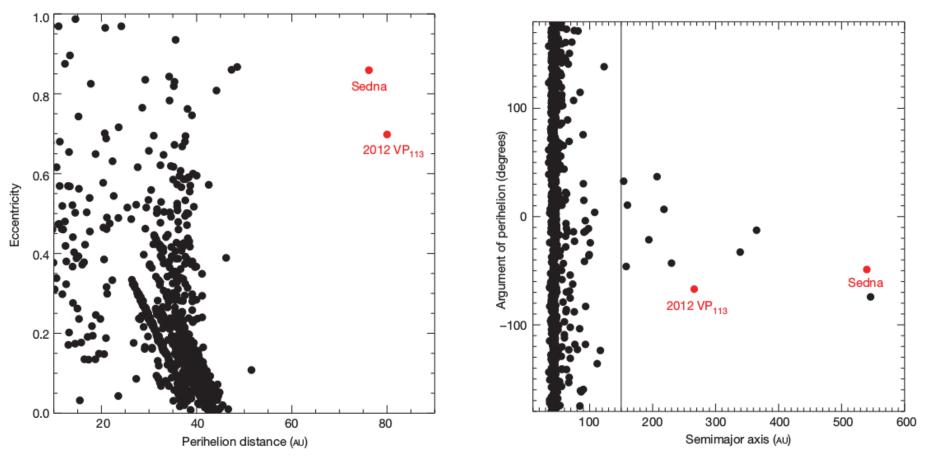
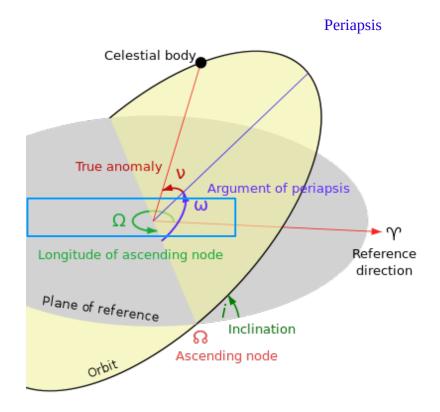


Figure 1 | Sedna and 2012 VP<sub>113</sub> are clear dynamical outliers in the Solar System. Eccentricity versus perihelion distance for the approximately 1,000 minor planets with well-determined (multi-year) orbits beyond 10 AU are

Figure 3 | The argument of perihelion for distant objects clusters about 0°. All minor planets with perihelion greater than 30 AU as a function of semimajor axis are shown. All bodies with semi-major axis greater than the line at 150 AU show a pronounced concentration near  $\omega \approx 0^\circ$ . Errors on these orbital

### $\rightarrow$ Consistent with simulated Super Earth-mass body at 250 AU

# Celestial mechanics: orbital elements (wikipedia)

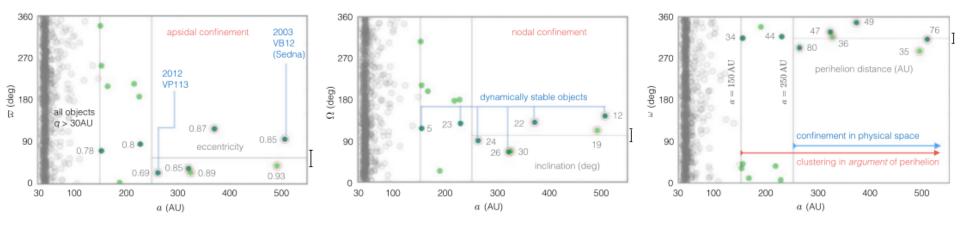


• Argument of periapsis  $\omega$ : angle from the body's ascending node to its periapsis, measured in the direction of motion *[Perihelion (or apsidal or orbital) precession*  $\omega$  = rotation of the orbit of a celestial body]

 $\omega_{\text{total}} = \omega_{\text{General Relativity}} + \omega_{\text{quadrupole}} + \omega_{\text{tide}} + \omega_{\text{perturbations}}$ 

- Longitude of ascending node  $\boldsymbol{\Omega}$
- Longitude of perihelion (or periapsis)  $\varpi = \omega + \Omega$

# Batygin & Brown, ApJ 151, 22 (2016)



#### **Observations**

- Confinement in  $\omega$  space
- Confinement in physical space

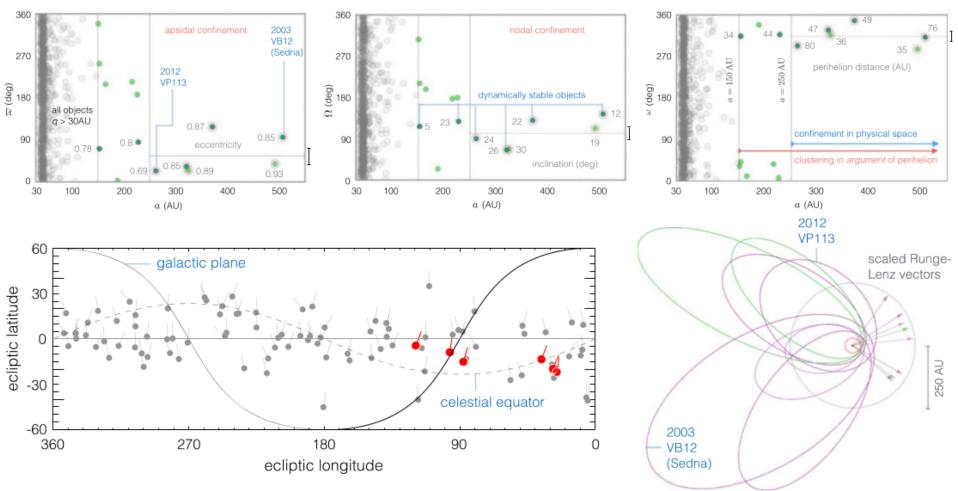
*N.B.*: strong observational bias towards w =0, but accounted for in analysis

• Argument of periapsis  $\omega$ : angle from the body's ascending node to its periapsis, measured in the direction of motion *[Perihelion (or apsidal or orbital) precession*  $\omega$  = rotation of the orbit of a celestial body]

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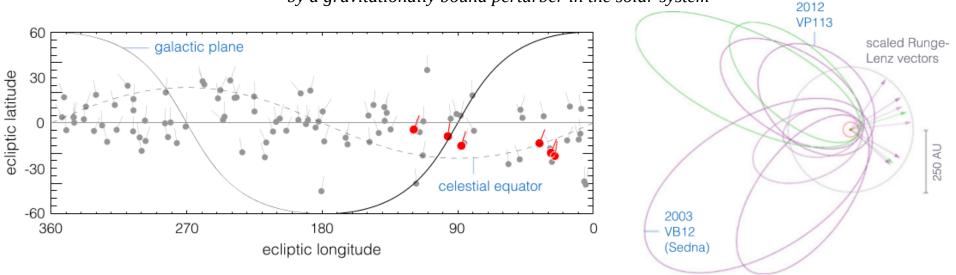
# Batygin & Brown, ApJ 151, 22 (2016)

**1)** Coherent dynamical structures in particle disks can be sustained by

- self-gravity (Tremaine 1998)
- gravitational shepherding by an extrinsic perturber (Goldreich & Tremaine 1982)

2) Mass of Kuiper Belt insufficient for self-gravity to play a role in its dynamical evolution

 $\rightarrow$  hypothesized that observed structure of the Kuiper Belt is maintained by a gravitationally bound perturber in the solar system



### Numerical simulations of this perturber

→ Although the model parameters are inherently degenerated, calculation suggests  $d\sim$ 700 AU,  $e\sim$ 0.6 (280-1120 AU), inclination=30 deg,  $M=10 \text{ m}_{\oplus}$ 

N.B.: long-term modulation of scattered KBO eccentricities provides a natural explanation for the existence of the so-called distant detached objects such as Sedna and 2012VP113

Discussions related to planet 9

### • Where is Telisto/Planet Nine? (Iorio, arXiv:1512.05288)

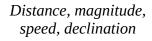
Use of Saturn ephemerides: modifications in secular precession of 
 *ω* [*N.B.: formalism used* ~ *to what MOND theory would give*]

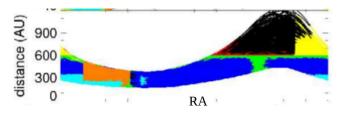
 $f \in [150-200] \text{ deg} \rightarrow a \in [930-1027] \text{ AU}$ 

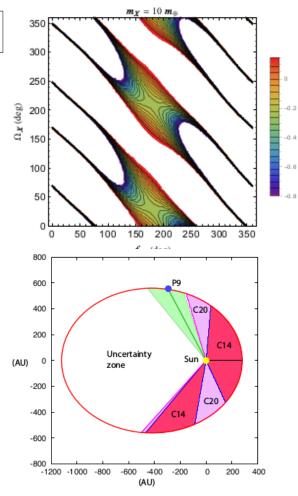
- New Horizons mission (~35 AU):  $\Delta x$ ~100 m by Telisto expected) [or dedicated mission to test gravitational laws?]
- Constraints from the Cassini data (Fienga et al., A&A 587, L8 (2016))
  - Also use all planets + Pluto + Moon ephemerides
  - Cross-check deviations with Cassini data
  - Global fit (150 params, 150000 observations)

Contradict Iori's conclusion: 'a' can be smaller than 1000 Motivation to extend Cassini up to 2020

- Observational constraints on the orbit (Brown & Batygin, arXiv:1603:05712)
  - WISE (µm), CRTS (Catalina Real-Time Transient Survey), DES (photometry), Pan-STARRS (PST: survey for transients), Pan-STARRS (moving object analysis), Pan-STARRS (expected limits), Ephemerides Saturne
  - Not excluded in previous or currently planned surveys







**Fig. 6.** Allowed zone for P9. The red zone (C14) is excluded by the analysis of the *Cassini* data up to 2014 (Sect. 4). The pink zone (C20) is how much this zone can be enlarged by extending the *Cassini* data to 2020 (Sect. 4). The green zone is the most probable zone for P9 ( $v \in [108^{\circ}:129^{\circ}]$ ), with maximum reduction of the residuals at  $v = 117.8^{\circ}$ 

Discussions related to planet 9

- Evolution and magnitude (Linder & Mordasini, arXiv:1602:07465)
  - Internal luminosity released may allow detection (isolated from Sun)
  - 50 M:, would have been seen
  - 10 M<sub>a</sub> (23.7 mag@1120 AU): OK with null detection

LSST (down to R=26 mag) or dedicated surveys should be able to find/exclude planet 9

- Blackbody rad. from isolated Neptunes (Ginzburg et al., arXiv:1603:02876)
  - Cooling study to relate surface T/luminosity to mass/composition
  - Use a two-layer Neptune model (rocky/icy core + H/He envelope)

$$T_{\rm eff} \approx 50 \ {\rm K} \left(\frac{R}{R_{\rm N}}\right)^{1/2} \left(\frac{M_{\rm atm}}{M_{\oplus}}\right)^{1/12} \left(\frac{t}{4.5 \ {\rm Gyr}}\right)^{-1/3} \qquad T_c \approx 2700 \ {\rm K} \left(\frac{M_{\rm atm}}{M_{\oplus}}\right)^{1/3}$$

#### Null detection can constrain the size and composition

• Cosmologist in search of planet 9 (Cowan et al.,arXiv:1602.05963)

Scaling Neptune: 
$$F_{1\text{mm}} = 30 \text{ mJy} \left(\frac{T_{\text{eff}}}{40 \text{ K}}\right) \left(\frac{R_p}{R_N}\right)^2 \left(\frac{d}{700 \text{ AU}}\right)^{-2}$$
,  
 $F_{2\text{mm}} = 8 \text{ mJy} \left(\frac{T_{\text{eff}}}{40 \text{ K}}\right) \left(\frac{R_p}{R_N}\right)^2 \left(\frac{d}{700 \text{ AU}}\right)^{-2}$ 

 Beware of false positives (stars, asteroids, Trans-Neptunians)
 Detection should be achieved with mm skymap every few months (30mJy, ~5 arcmn)

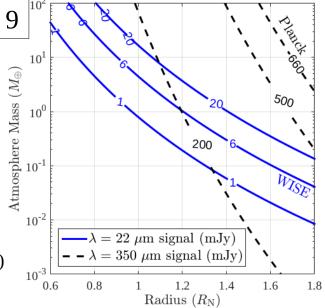
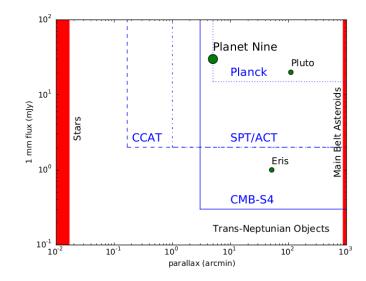


FIG. 4.— Contours of the observable flux density at wavelengths of 22  $\mu$ m (solid blue lines) and 350  $\mu$ m (dashed black lines) of 4.5 Gyr old planets at a distance of 700 AU, as a function of their radius and atmosphere mass. The 22  $\mu$ m sensitivity of WISE is 6 mJy (Wright et al. 2010), corresponding to the middle solid blue line, while the 350  $\mu$ m sensitivity of Planck is 660 mJy (Planck Collaboration et al. 2014, 2015), corresponding to the top dashed black line.



Discussions related to planet 9

### • Interaction Cross Sections and Survival Rates (Li & Adams, arXiv:1602:08496)

Astronomy has a long history of considering gravitational perturbations to uncover new Solar system bodies

- Neptune (LeVerrier 1846; Galle 1846; Adams 1846)
- Large Planet X beyond Neptune: never found, but searches led to Pluto (see Tombaugh 1996)
- Nemesis (red dwarf companion to the Sun) to explain periodic extinctions in fossil record...

*This paper*: possible formation scenarios and dynamical histories for Planet Nine (wide orbit, so highly susceptible to gravitational perturbations from passing stars and binaries).

Formation scenario:

- formed in a nearly circular orbit with a much smaller semimajor axis (a a 9  $\sim$  400 1500 AU)
- scattered by the giant planets into an orbit with perihelion p  $\sim$  30 40 AU
- captured from another solar system, and/or as a freely-floating planet within the Solar birth cluster
- can be removed from the Solar System by passing stars

Sun Environment:

- Sun formed in cluster with  $10^3 10^4$  stars + stay for ~ 100 Myr in  $< n_* > \sim 100 \text{ pc}^{-3}$  and  $v_h \sim 1 \text{ km/s}$
- After leaving cluster, SS ~4.5 Gyr in Solar neighborhood with  $\langle n_* \rangle \sim 1 pc^{-3}$ ,  $v_b \sim 40 km/s$

Simulations result:

Scattering simulations indicate that survival of Planet Nine is possible but not guaranteed Formation scenario for Planet Nine represents an important challenge for future investigation