

## *Nuclear PDFs*

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# OUTLINE

- ① INTRODUCTION
- ② REVIEW OF GLOBAL ANALYSES OF NUCLEAR PDFS
- ③ NUCLEAR EFFECTS IN  $\nu A$  DIS
- ④ THE NUCLEAR GLUON DISTRIBUTION

# Introduction

# NUCLEAR PDFs (NPDF)

- Information on **hadronic structure**
- **Initial state** for hard processes in collisions involving hadrons
  - Deep inelastic scattering (DIS):  $\ell A, \nu A$
  - Drell-Yan (DY):  $A + B \rightarrow \ell^+ + \ell^-$
  - Jets, Photons, Hadrons at large  $p_T$ ; Heavy Quarks; ...  
in  $pA, AA, (\gamma A, eA)$  collisions
- Provide **nuclear corrections** for global analyses of **proton PDFs** in a **flexible way**

# THEORETICAL BASIS: FACTORIZATION

- Factorization theorems
  - provide (field theoretical) **definitions of universal PDFs**
  - make the formalism **predictive**
  - make a statement about the **error**
- **PDFs and predictions for observables+uncertainties refer to this standard pQCD framework**
- There might be breaking of QCD factorization, deviations from **DGLAP** evolution — in particular in a nuclear environment

**Still need solid understanding of standard framework  
to establish deviations!**

In the nuclear case, consider factorization as a **working assumption** to be tested phenomenologically

# Review of global analyses of nuclear PDFs

# NPDFs FROM $\ell A$ DIS AND DY DATA

Global analyses of NPDF by four groups:

- **HKN'07** [PRC76(2007)065207]  
LO, NLO, error PDFs,  $\chi^2/dof = 1.2$
- **EPS'09** [JHEP0904(2009)065]  
LO, NLO, error PDFs,  $\chi^2/dof = 0.8$   
Use also inclusive  $\pi^0$  data at midrap. from  
 $d + Au$  and  $p + p$  coll. at RHIC  $\rightarrow$  gluon
- **DS'04** [PRD69(2004)074028]  
first NLO analysis, 'semi-global', no error  
PDFs,  $\chi^2/dof = 0.76$
- **nCTEQ** [PRD80(2009)094004]  
NLO, same data as HKN'07 (up to cuts),  
no error PDFs (so far),  $\chi^2/dof = 0.95$ ,  
official release soon

Table from Hirai et al., arXiv:0909.2329

R	Nucleus	Experiment	EPS09	HKN07	DS04
DIS	D/p	NMC		O	
	4He	SLAC E139	O	O	O
	Li	NMC95	O (5)	O	O
	Be	SLAC E139	O	O	O
	C	EMC-88, 90 NMC 95 SLAC E139 FNAL-E665		O O O	
	N	BCDMS 85 HERMES 03		O	
	Al	SLAC E49 SLAC E139		O O	O
	Ca	EMC 90 NMC 95 SLAC E139 FNAL-E665		O O O	
	Fe	SLAC E87 SLAC E139 SLAC E140 BCDMS 87	O (15)	O	O
	Cu	EMC 93	O	O	
	Kr	HERMES 03		O	
	Ag	SLAC E139	O	O	O
	Sn	EMC 88		O	
	Au	SLAC E139 SLAC E140	O O	O	O
	Pb	FNAL-E665		O	
	Be	NMC 96	O	O	O
	Al	NMC 96	O	O	O
	Ca	NMC 95		O	
	Fe	NMC 96	O	O	O
	Sn	NMC 96	O (10)	O	O
	Pb	NMC 96	O	O	O
DY	A/Li	C Ca	O O	O	O
	C	FNAL-E772	O (15)	O	O
	Ca		O (15)	O	O
	Fe		O (10)	O	O
	W	FNAL E866	O	O	
	A/Be		O	O	
$\pi$ pro	W		O	O	
	dA/pp	Au	RHIC-PHENIX	O (20)	

# WHAT ARE THE DIFFERENCES?

## Main differences:

- **Choice of data sets** (see previous table)
  - **Parametrization of input distributions**
  - **Assumptions on PDFs**
    - Data less constraining than in proton case → need to make more assumptions (otherwise flat directions in  $\chi^2$  function and fits don't converge)
    - Assumptions replace uncertainty! → error bands (of a single fit) underestimate true uncertainties
- 

## Consequences?

- **Use different sets of NPDFs to scan over assumptions**
- Include more data sets → allows to relax assumptions
- New ideas to handle flat directions?
- Neural Network PDFs?

# WHAT ARE THE DIFFERENCES?

Further differences:

- **Heavy flavor schemes**

- **DS'04:** 3-Fixed Flavor Number Scheme (3-FFNS) → no charm PDF
- **HKN'07, EPS'09, nCTEQ:** Variable Flavor Number Schemes (VFNS)

→ Beware of comparing 'apples with oranges'!

- **Parameters and other**

- Input scale  $Q_0$ ,  $\alpha_s(M_Z)$ ,  $m_c$ ,  $m_b$
- Evolution in  $n$ -space (DS) and  $x$ -space (HKN,EPS,nCTEQ)
- Target Mass Corrections (TMC)  
see, e.g., [IS et al., JPG35(2008)053101; Qiu, Accardi, JHEP0807(2008)090]

Connected to GRV'98 proton PDFs  $f_i^p(x, Q)$ :

- $Q_0^2 = 0.4 \text{ GeV}^2$  (NLO),  $Q_0^2 = 0.26 \text{ GeV}^2$  (LO),  $m_c$ ,  $m_b$ ,  $\alpha_s$  as in GRV'98
- 3-Fixed flavor scheme (no charm PDF)
- strange PDF dynamically generated, i.e.,  $s^p(x, Q_0^2) = 0$

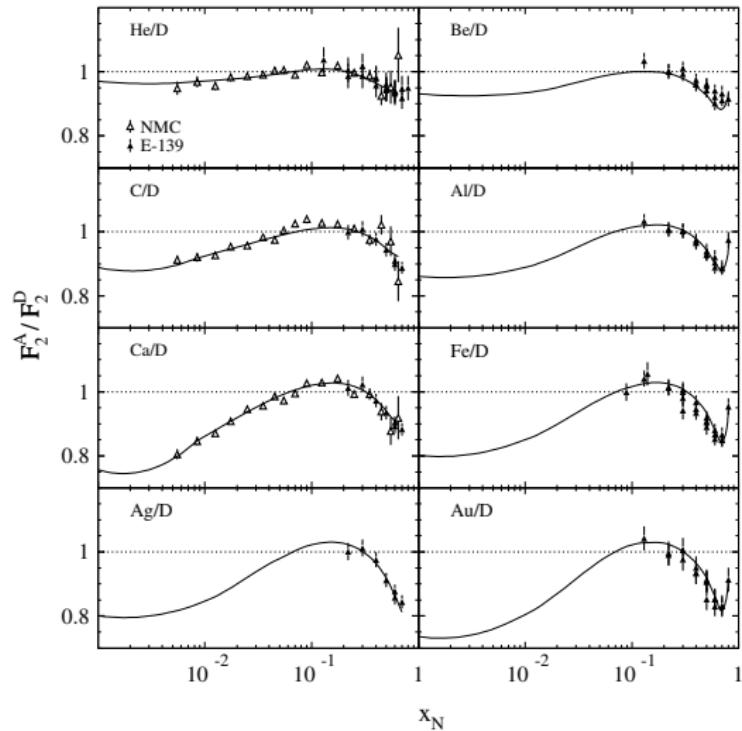
Parametrization of input distributions:

- PDFs for bound protons inside nucleus A:  $f_i^{p/A}(x, Q)$
- Convolution relation: 
$$f_i^{p/A}(x_N, Q_0^2) = \int_{x_N}^A \frac{dy}{y} W_i(y, A, Z) f_i^p(x_N/y, Q_0^2)$$
- Weight functions  $W_v$  (valence),  $W_s$  (sea),  $W_g$  (gluon). For example:

$$\begin{aligned} W_v(y, A, Z) &= A[a_v \delta(1 - \epsilon_v - y) + (1 - a_v) \delta(1 - \epsilon_{v'} - y)] \\ &\quad + n_v (y/A)^{\alpha_v} (1 - y/A)^{\beta_v} + n_s (y/A)^{\alpha_s} (1 - y/A)^{\beta_s} \end{aligned}$$

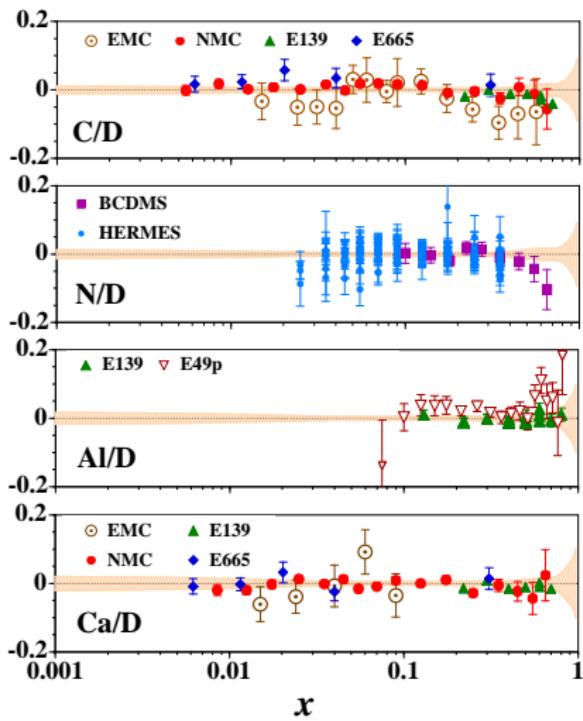
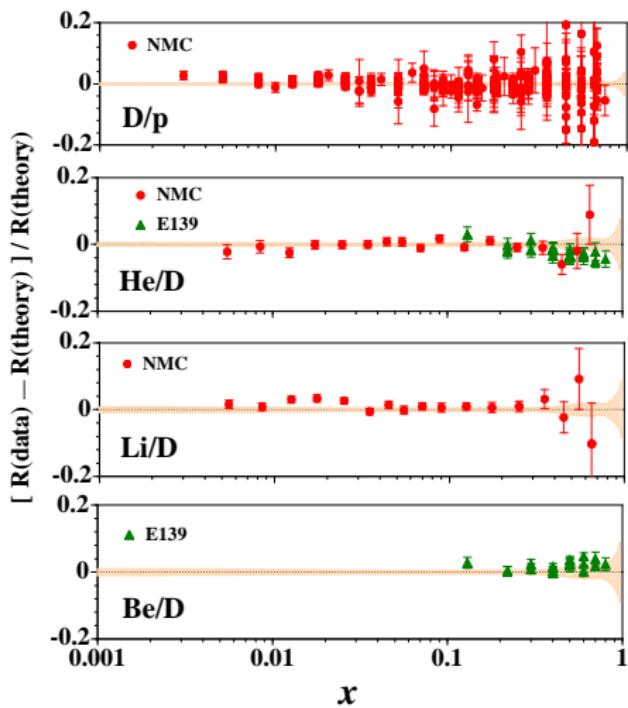
- Note:
  - Convolution simple product in Mellin moment space: very elegant
  - Ansatz valid for  $0 < x_N < A$ !
  - The  $x$ -space approaches (HKN, EPS, nCTEQ) are restricted to  $0 < x_N < 1$
  - However, the DS'04 PDF grids apparently are restricted to  $0 < x_N < 1$  (and the momentum sum rule integrates to unity in this range)

Excellent fit to a **restricted** data set (420 points):  $\chi^2/\text{dof} = 0.75$



- LO, NLO, error PDFs
- Related to MRST'98 proton PDF:  $Q_0^2 = 1 \text{ GeV}^2$
- Uses multiplicative ansatz:  $f_i^{p/A}(x_N, Q_0^2) = R_i(x_N, Q_0, A, Z) f_i^p(x_N, Q_0^2)$
- Weight factor:  $R_i(x, A, Z) = 1 + (1 - \frac{1}{A^\alpha}) \frac{a_i + b_i x + c_i x^2 + d_i x^3}{(1-x)^{\beta_i}}$  ( $i = u_v, d_v, \bar{q}, g$ )
- neglects region  $x_N > 1$
- includes all current DIS & DY data sets, in particular deuterium data
- uses Hessian method to produce error PDFs

- Reasonable fits:  $\chi^2/dof = 1.2$



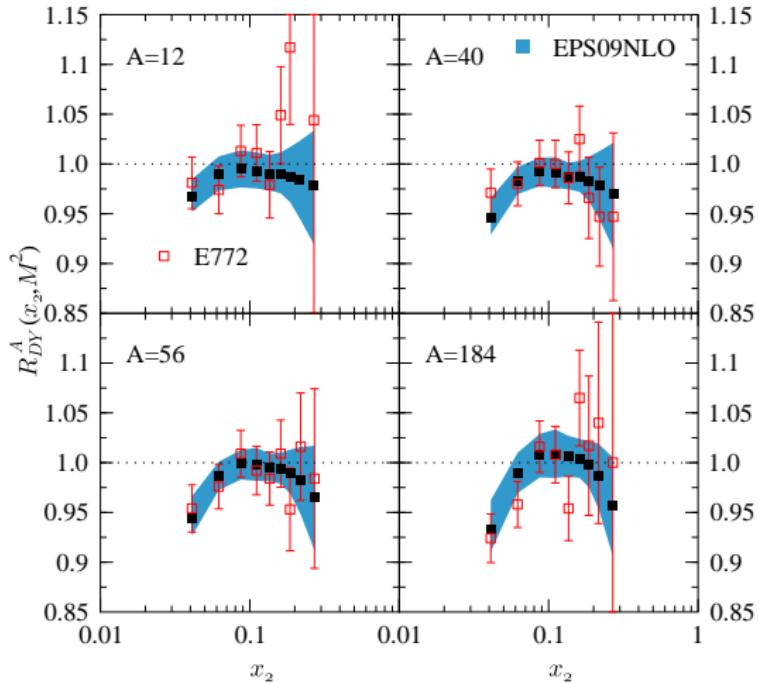
- LO, NLO, error PDFs
- Related to CTEQ6.1M proton PDF:  $Q_0 = 1.3 \text{ GeV}$
- Uses multiplicative ansatz:  $f_i^{p/A}(x_N, Q_0^2) = R_i(x_N, Q_0, A, Z) f_i^p(x_N, Q_0^2)$
- Weight factor is a piecewise defined function:

$$R_i(x, A, Z) = \begin{cases} a_0 + (a_1 + a_2 x)(e^{-x} - e^{-x_a}) & x \leq x_a \\ b_0 + b_1 x + b_2 x^2 + b_3 x^3 & x_a \leq x \leq x_e \\ c_0 + (c_1 - c_2 x)(1 - x)^{-\beta} & x_e \leq x \leq 1 \end{cases}$$

where the parameters  $a_i, b_i, c_i, \beta, x_a, x_e$  are  $A$ -dependent

- neglects region  $x_N > 1$
- includes  $\pi^0$  RHIC data with a weight 20 to constrain gluon
- uses Hessian method to produce error PDFs

- Excellent fit:  $\chi^2/dof = 0.8$
- Show here, as an example, comparison with DY data



# NUCLEAR CTEQ

## Work in collaboration with:

- People from LPSC Grenoble: K. Kovarik, J. Y. Yu, T. Stavreva, IS
  - CTEQ-members: F. Olness (SMU), J. Owens (FSU), J. Morfin (FNAL), C. Keppel (JLAB)
- 

- The results shown in the following are from  
IS,Yu,Kovarik,Keppel,Morfin,Olness,Owens,PRD80(2009)094004
- A first set of nCTEQ nuclear PDFs will be released in the near future

# NUCLEAR CTEQ

Framework as in CTEQ6M proton fit:

- **Same functional form** for **bound proton PDFs** inside a nucleus  $A$  as for free proton PDFs (restrict  $x$  to  $0 < x < 1$ ):

$$\begin{aligned} x f_k^{p/A}(x, Q_0) &= c_0 x^{c_1} (1-x)^{c_2} e^{c_3 x} (1 + e^{c_4 x})^{c_5}, \quad k = u_V, d_V, g, \bar{u} + \bar{d}, s, \bar{s}, \\ \bar{d}(x, Q_0)/\bar{u}(x, Q_0) &= c_0 x^{c_1} (1-x)^{c_2} + (1 + c_3 x)(1-x)^{c_4} \end{aligned}$$

(bound neutron PDFs  $f_k^{n/A}$  by isospin symmetry)

- **$A$ -dependent fit parameters:** (reduces to free proton parameters  $c_{k,0}$  for  $A = 1$ )

$$c_k \rightarrow c_k(A) \equiv c_{k,0} + c_{k,1}(1 - A^{-c_{k,2}}), \quad k = 1, \dots, 5$$

- **PDFs for a nucleus** ( $A, Z$ ):  $f_i^{(A,Z)}(x, Q) = \frac{Z}{A} f_i^{p/A}(x, Q) + \frac{A-Z}{A} f_i^{n/A}(x, Q)$
- **Input parameters:**  $Q_0 = m_c = 1.3$  GeV,  $m_b = 4.5$  GeV,  $\alpha_s^{NLO, \overline{\text{MS}}}(M_Z) = 0.118$
- **Heavy quark treatment:** ACOT scheme
- **Standard DIS-cuts:**  $Q > 2$  GeV,  $W > 3.5$  GeV

# EXPERIMENTAL INPUT

Use same data as HKN'07 (up to cuts)

- DIS  $F_2^A/F_2^D$  data sets: 862 points (before cuts)
- DIS  $F_2^A/F_2^{A'}$  data sets: 297 points (before cuts)
- DY data sets  $\sigma_{\text{DY}}^{pA}/\sigma_{\text{DY}}^{pA'}$ : 92 points (before cuts)

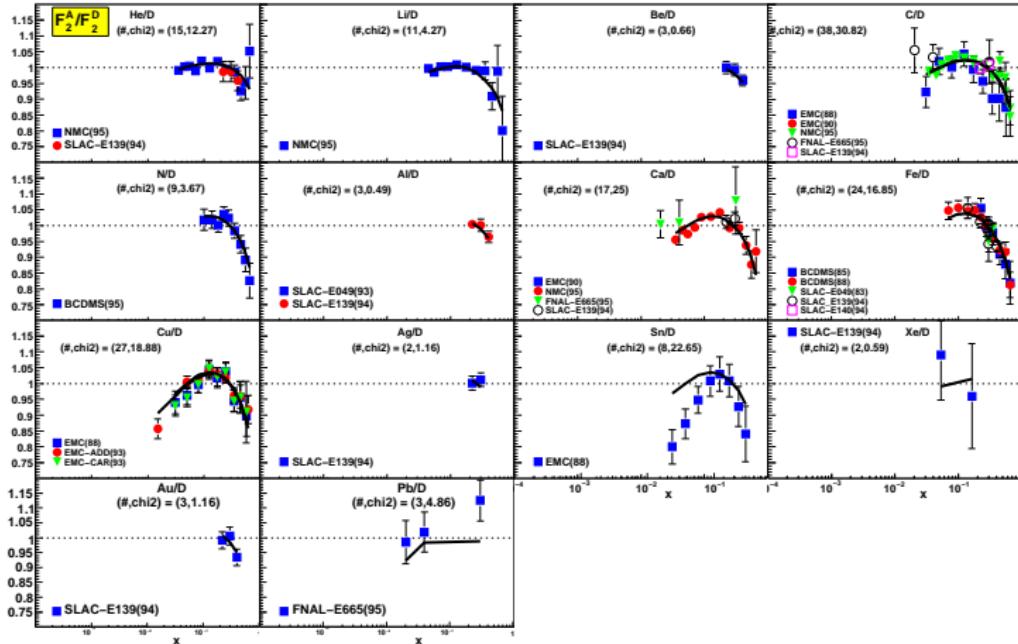
Table from Hirai et al., arXiv:0909.2329

R	Nucleus	Experiment	EPS09	HKN07	DS04
DIS	D/p	NMC		O	
	4He	SLAC E139	O	O	O
		NMC95	O (5)	O	O
	Li	NMC95	O	O	
	Be	SLAC E139	O	O	O
		EMC-88, 90		O	
	C	NMC 95	O	O	O
		SLAC E139	O	O	O
		FNAL-E665		O	
	N	BCDMS 85		O	
		HERMES 03		O	
	Al	SLAC E49		O	
		SLAC E139	O	O	O
		EMC 90		O	
	Ca	NMC 95	O	O	O
		SLAC E139	O	O	O
		FNAL-E665		O	
		SLAC E87		O	
A/C	Fe	SLAC E139	O (15)	O	O
		SLAC E140		O	
		BCDMS 87		O	
	Cu	EMC 93	O	O	
	Kr	HERMES 03		O	
	Ag	SLAC E139	O	O	O
	Sn	EMC 88		O	
	Au	SLAC E139	O	O	O
		SLAC E140		O	
	Pb	FNAL-E665		O	
DY	Be	NMC 96	O	O	O
	Al	NMC 96	O	O	O
	Ca	NMC 95		O	
	Fe	NMC 96	O	O	O
	Sn	NMC 96	O (10)	O	O
	Pb	NMC 96	O	O	O
π pro	C	NMC 95	O	O	
	Ca	NMC 95	O	O	
	C		O	O	O
	Ca	FNAL-E772	O (15)	O	O
	Fe		O (15)	O	O
dA/dp	W		O (10)	O	O
	A/Be	FNAL E866	O	O	
	W		O	O	
	dA/dp	Au RHIC-PHENIX	O (20)		

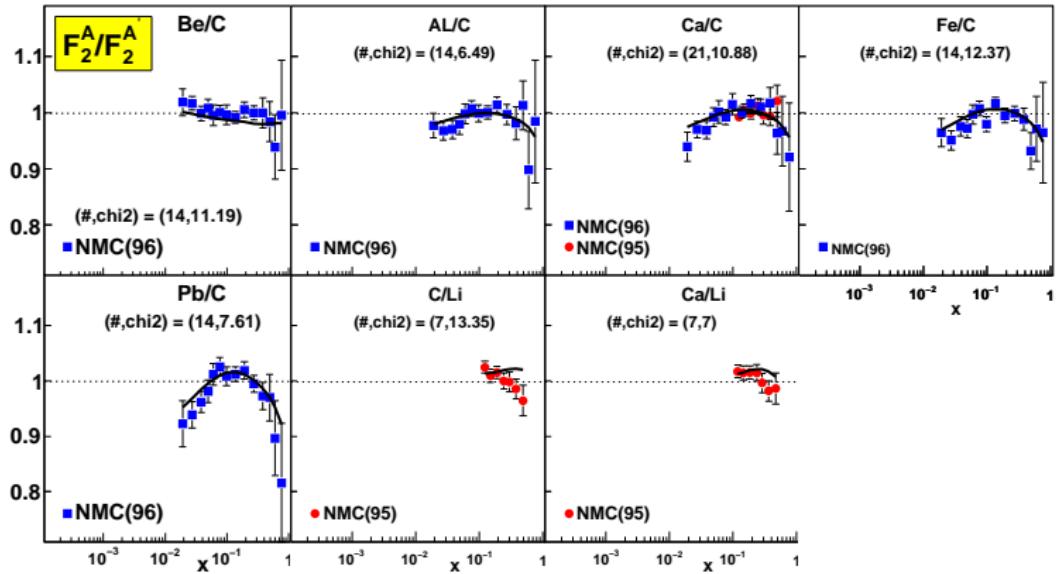
## RESULTS: DECUT3 FIT

- 708 (1233) data points after (before) cuts
- 32 free parameters; 675 d.o.f.
- Overall  $\chi^2/\text{d.o.f.} = 0.95$
- individually:
  - for  $F_2^A/F_2^D$ :  $\chi^2/\text{pt} = 0.92$
  - for  $F_2^A/F_2^{A'}$ :  $\chi^2/\text{pt} = 0.69$
  - for DY:  $\chi^2/\text{pt} = 1.08$
- **Our simple approach works!**

# RESULTS: DECUT3 FIT DIS DATA VS $x$

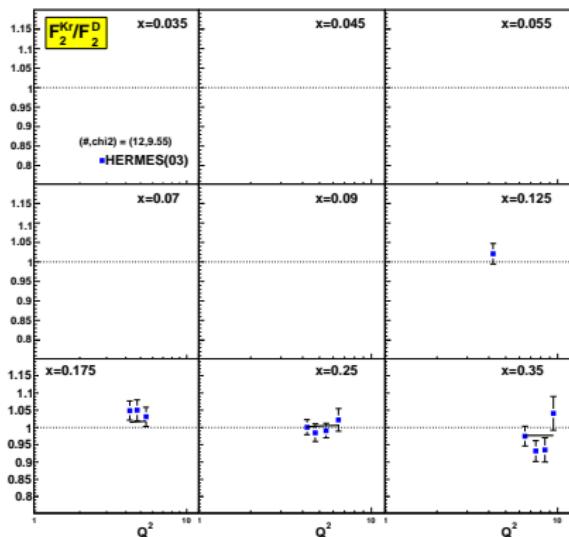
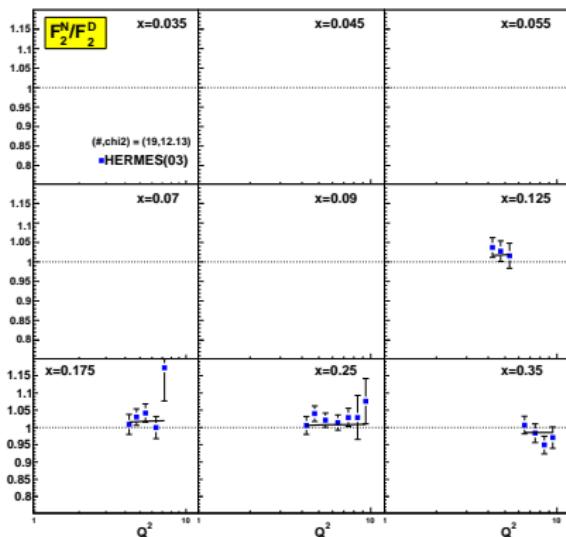


# RESULTS: DECUT3 FIT DIS DATA VS $x$



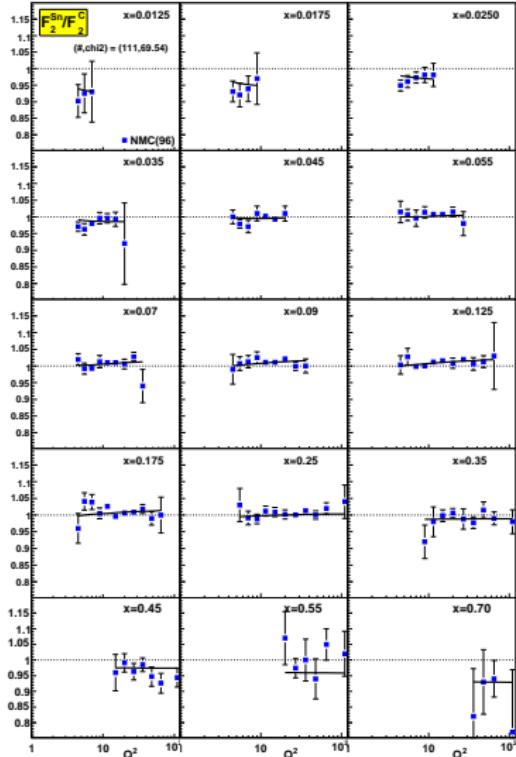
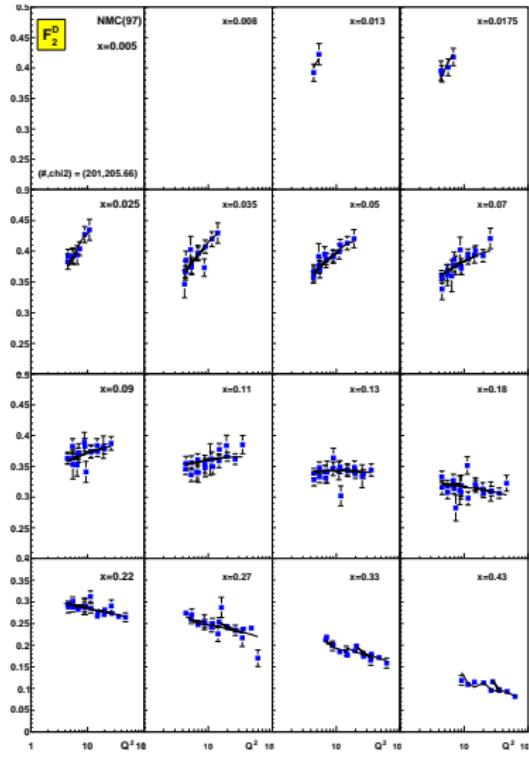
# RESULTS: DECUT3 FIT

## HERMES DATA VS $Q^2$



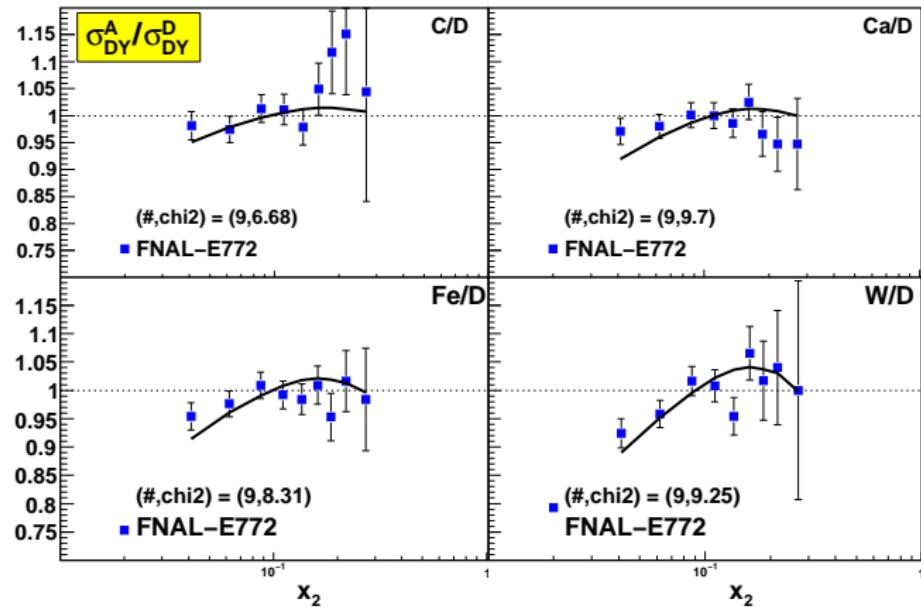
# RESULTS: DECUT3 FIT

## NMC DATA FOR $D$ AND $Sn/C$ vs $Q^2$



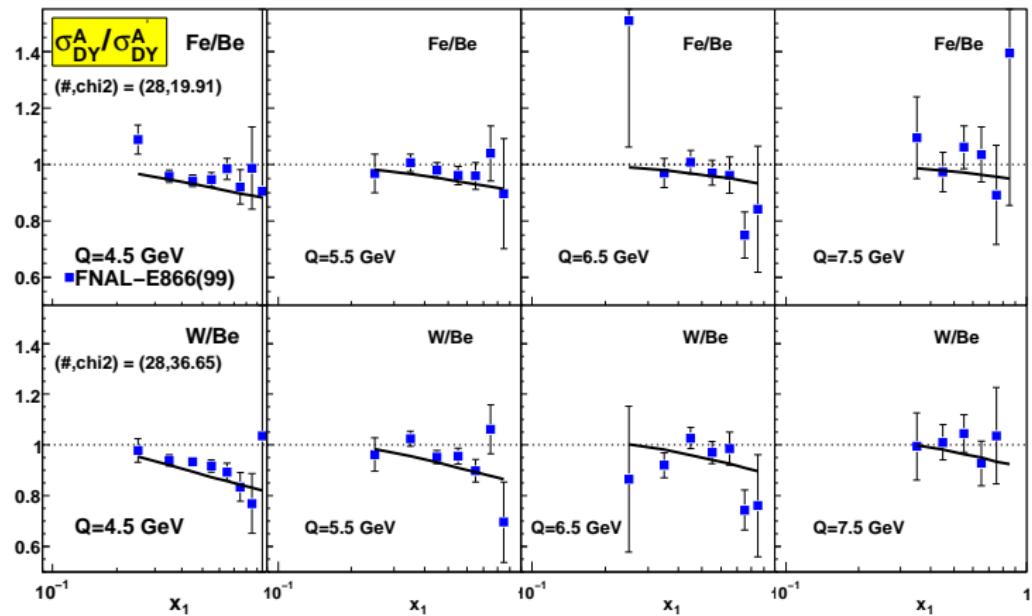
# RESULTS: DECUT3 FIT

DRELL-YAN DATA



# RESULTS: DECUT3 FIT

DRELL-YAN DATA



# CONCLUSIONS I

- Factorization works well for  $\ell A$  DIS and DY data
- Nuclear CTEQ PDFs will be released in the near future

# Nuclear effects in $\nu A$ DIS

I.S., Yu, Keppel, Morfin, Olness, Owens, PRD77(2008)054013

I.S., Yu, Kovarik, Keppel, Morfin, Olness, Owens, PRD80(2009)094004

Kovarik, Yu, Keppel, Morfin, Olness, Owens, I.S., Stavreva, work nearing completion

# WHY NEUTRINO DIS?

- **Flavor separation:**

Neutrino sfs depend on different combinations of PDFs

- **Dimuon production:**

- Main source of information on the strange sea
- Large uncertainty on  $s(x, Q^2)$  has significant influence on the  $W$  and  $Z$  benchmark processes at the LHC

- Data interesting for proton PDF and NPDF

- For proton PDF: need **nuclear corrections**

- **EW precision measurements:**

Paschos-Wolfenstein analysis: extraction of  $\sin^2 \theta_W$

- **LBL precision neutrino experiments:**

Need good understanding of neutrino–nucleus cross sections

# NUCLEAR CORRECTION FACTORS

Be  $\mathcal{O}$  an **observable** calculable in the parton model

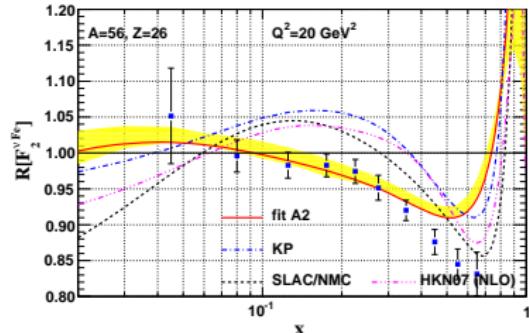
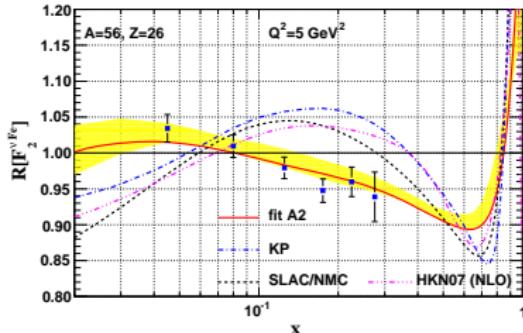
Define **nuclear correction factor**:

$$R[\mathcal{O}] := \frac{\mathcal{O}[\text{NPDF}]}{\mathcal{O}[\text{PDF}]} \quad \text{or for data} \quad R[\mathcal{O}] := \frac{\mathcal{O}^{\text{exp}}}{\mathcal{O}[\text{PDF}]}$$

- Factor needed to correct data to the free nucleon level
- Note: different observables  $\Rightarrow$  different correction factors
- In particular, correction factor for  $F_3^{\nu A}$  could be quite different from  $F_2^{\nu A}$ !
- Also  $R[F_2^{\ell A}]$ ,  $R[F_2^{\nu A}]$ ,  $R[F_2^{\bar{\nu} A}]$ ,  $R[d^2\sigma^{\nu A}/dx dQ^2]$ , ... are all (more or less) different **even for universal nPDFs**

Note: the term “nuclear effects” is less precise and (mis-)used in the literature for a lot of different things

# NUCLEAR CORRECTION FACTOR $R[F_2^{\nu Fe}]$



- Are nuclear corrections in charged-lepton and neutrino DIS different?
- Obviously the PDFs from fits to  $\ell A + \text{DY}$  data do not describe the NuTeV  $\nu Fe$  DIS data.
- A global analysis of  $\ell A + \text{DY} + \nu A$  data confirms this result!  
(see backup slides for a detailed account)

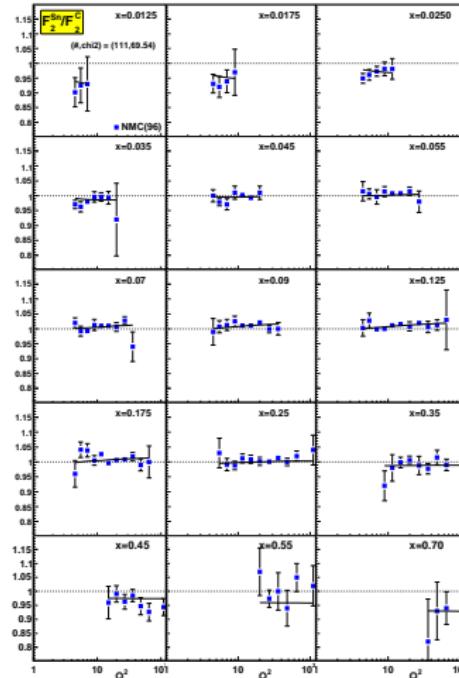
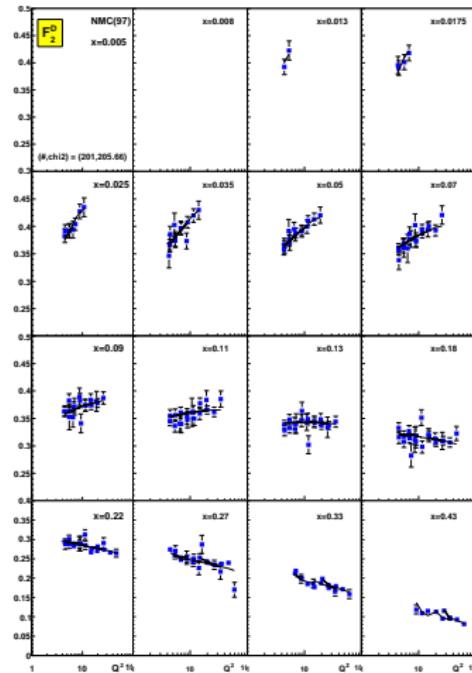
Note:  $x_{\min} = 0.02$  in these figures.

## CONCLUSIONS II

- We observe different nuclear effects in  $\ell A + DY$  data as opposed to NuTeV  $\nu Fe$  data
- These are precision effects relevant for precision observables
- Paukkunen and Salgado come to different conclusions in a recent paper
- The main reason for the different results is that Paukkunen and Salgado use uncorrelated systematic errors, whereas we take into account the full error correlation matrix
- For more details see backup slides and a publication in preparation

# The nuclear gluon distribution

# $g^A(x, Q^2)$ WEAKLY CONSTRAINED BY $Q^2$ -DEPENDENCE OF NMC DATA

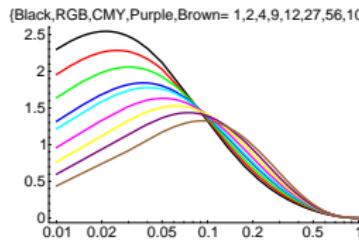


- $x \sim 0.01 \dots 0.4$ ,  $Q^2 \sim 10 \dots 100$  GeV $^2$

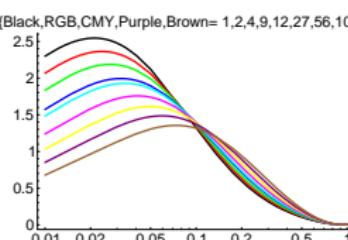
# THE NUCLEAR GLUON DISTRIBUTION

A series of **equally good fits** ( $\chi^2/pt \simeq 0.9$ ) to  $\ell A + \text{DY}$  data with different gluons

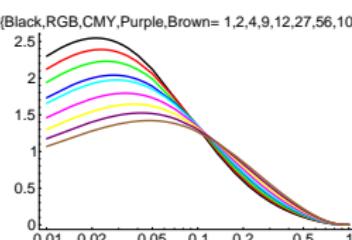
decut3



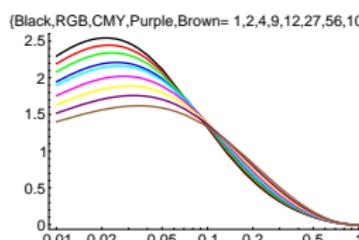
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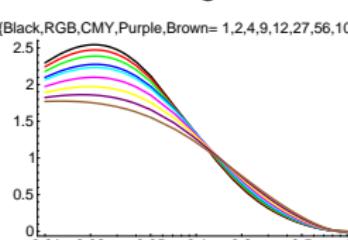
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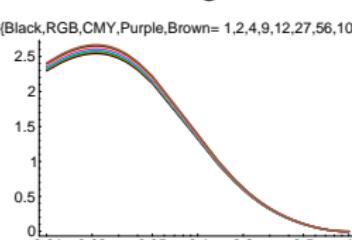
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decut3g8



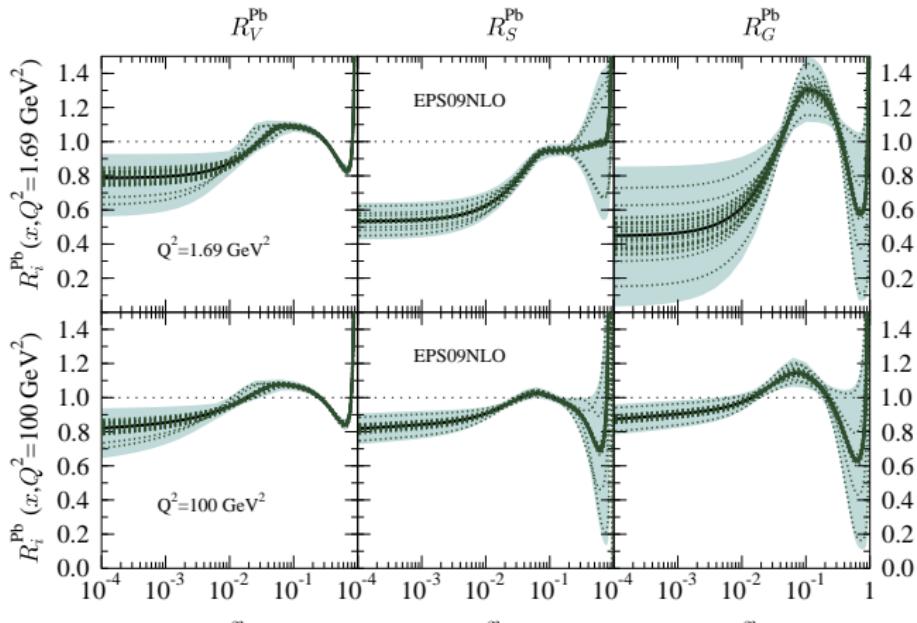
decut3g9



Shown are the gluon distributions at the scale  $Q_0 = 1.3 \text{ GeV}$  for different  $A$  vs  $x$

# GLUON UNCERTAINTY IN EPS'09

- EPS'09 also uses RHIC data for inclusive pion production to constrain the gluon
- This involves fragmentation functions  $D_i^\pi(z, \mu^2)$  into pions
- Large uncertainties! Still some of the gluons of the decut3g series lie outside the error band of  $R_G^{Pb}$



# NEED HARD PROBES IN $pA$ TO CONSTRAIN NPDFs

Hard probes in  $pp$ ,  $p\bar{p}$  to constrain proton PDFs:

- Tevatron inclusive jet data → gluon
- Lepton pair production → sea quarks
- Vector boson production → sea quarks (less useful due to high scale)

Other interesting processes:

- Prompt photon production [see Arleo, Gousset] → gluon
- Heavy quark production → gluon?
- $\gamma + j \rightarrow$  gluon
- $\gamma + j_Q$  (see talk by T. Stavreva) → gluon, charm
- $\gamma + J/\Psi$  (see talk by M. Machado)
- Quarkonium production?

# Backup slides

# Analysis of $\nu A$ , $\ell A$ and DY data

Kovarik,Yu,Keppel,Morfin,Olness,Owens,Schienbein,Stavreva, work nearing completion

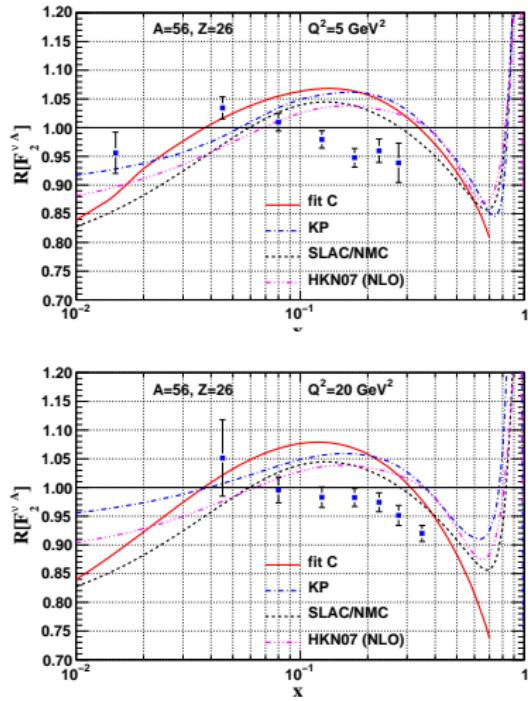
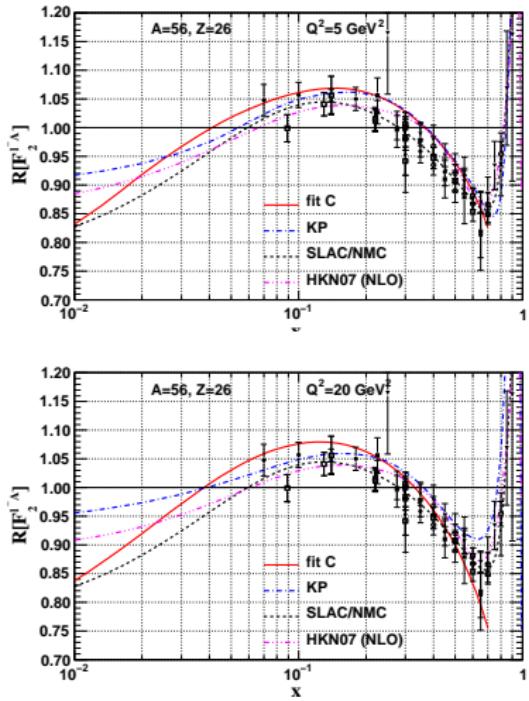
# COMBINING $\ell A$ DIS, DY AND $\nu A$ DIS DATA

- $\ell A$  and DY data sets as before
- 8 Neutrino data sets
  - NuTeV cross section data:  $\nu Fe$ ,  $\bar{\nu} Fe$
  - CHORUS cross section data:  $\nu Pb$ ,  $\bar{\nu} Pb$
  - NuTeV dimuon data:  $\nu Fe$ ,  $\bar{\nu} Fe$
  - CCFR dimuon data:  $\nu Fe$ ,  $\bar{\nu} Fe$
- Problem: Neutrino data sets have much higher statistics. Systematically study fits with different weights.

Weight	Fit name	$\ell$ data	$\chi^2$ (/pt)	$\nu$ data	$\chi^2$ (/pt)	total $\chi^2$ (/pt)
$w = 0$	decut3	708	639 (0.90)	-	-	639 (0.90)
$w = 1/7$	glofac1a	708	645 (0.91)	3134	4710 (1.50)	5355 (1.39)
$w = 1/4$	glofac1c	708	654 (0.92)	3134	4501 (1.43)	5155 (1.34)
$w = 1/2$	glofac1b	708	680 (0.96)	3134	4405 (1.40)	5085 (1.32)
$w = 1$	global2b	708	736 (1.04)	3134	4277 (1.36)	5014 (1.30)
$w = \infty$	nuanua1	-	-	3134	4192 (1.33)	4192 (1.33)

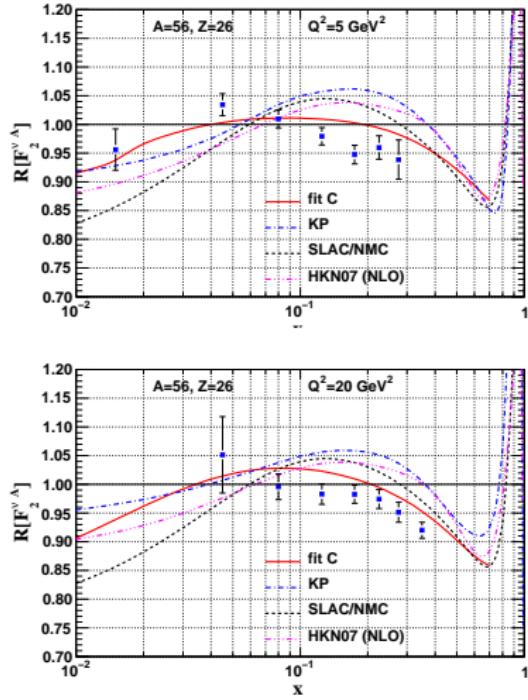
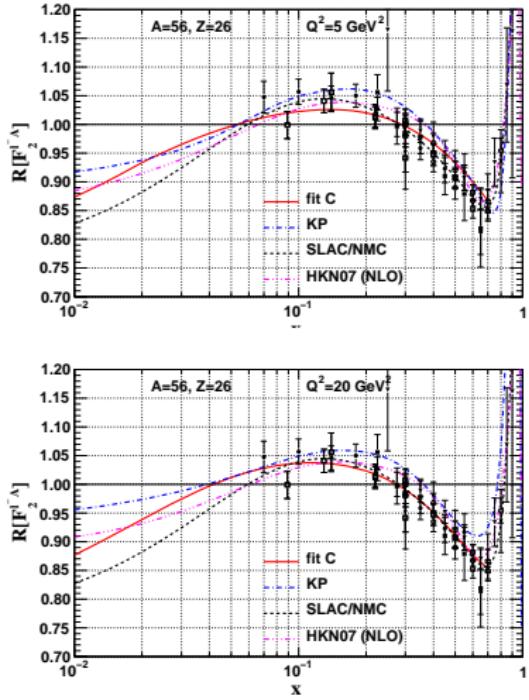
# $R[F_2^{\ell Fe}]$ (LEFT) VS $R[F_2^{\nu Fe}]$ (RIGHT)

decut3 ( $w = 0$ )



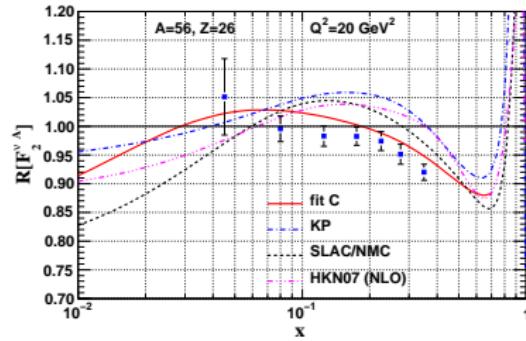
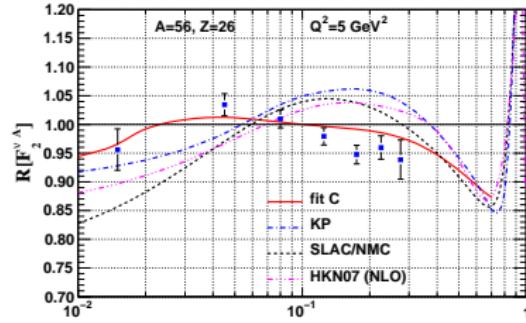
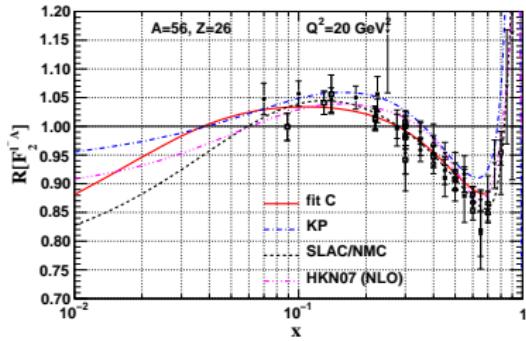
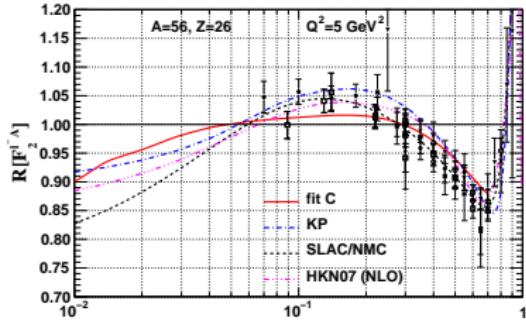
# $R[F_2^{\ell Fe}]$ (LEFT) VS $R[F_2^{\nu Fe}]$ (RIGHT)

glofac1a ( $w = 1/7$ )



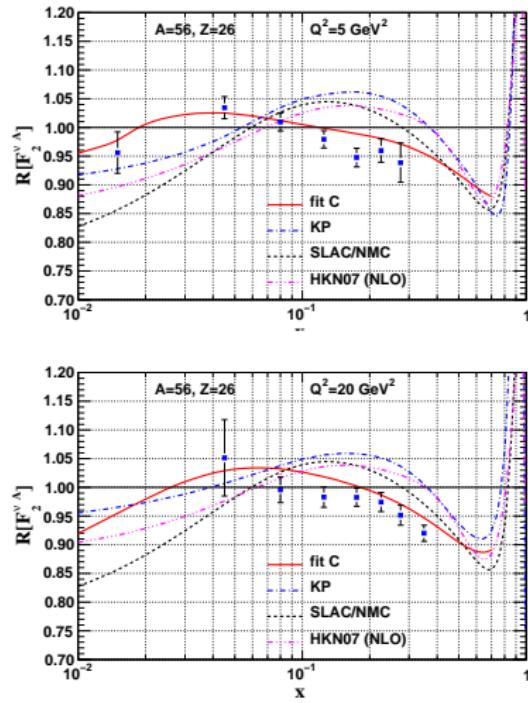
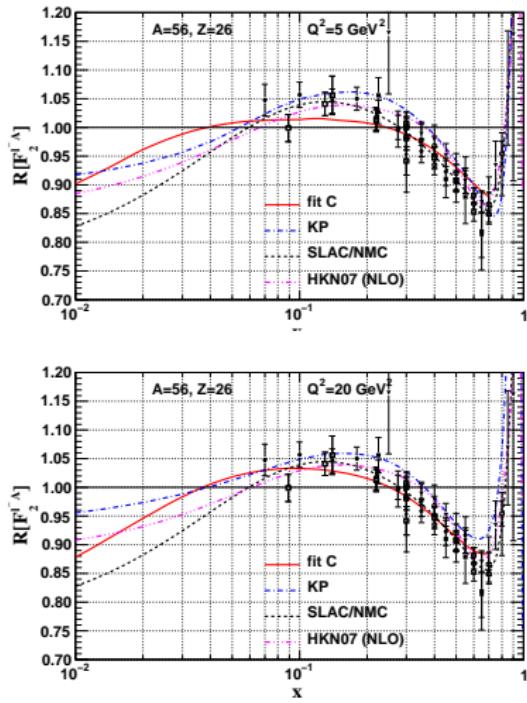
# $R[F_2^{\ell Fe}]$ (LEFT) VS $R[F_2^{\nu Fe}]$ (RIGHT)

glofac1c ( $w = 1/4$ )



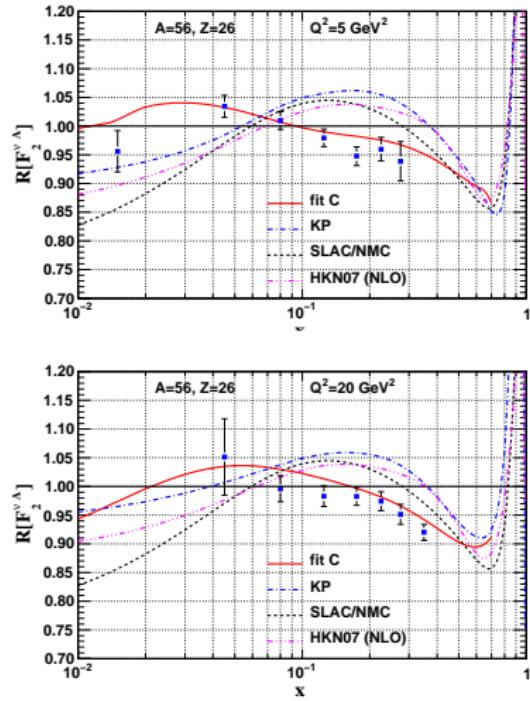
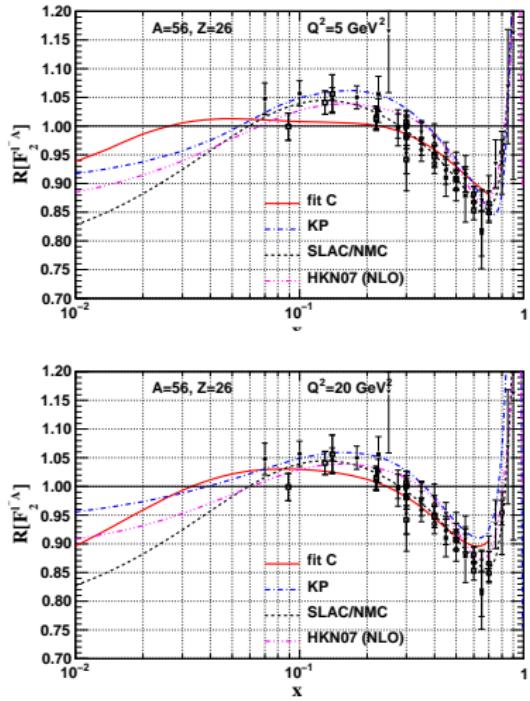
# $R[F_2^{\ell Fe}]$ (LEFT) VS $R[F_2^{\nu Fe}]$ (RIGHT)

glofac1b ( $w = 1/2$ )



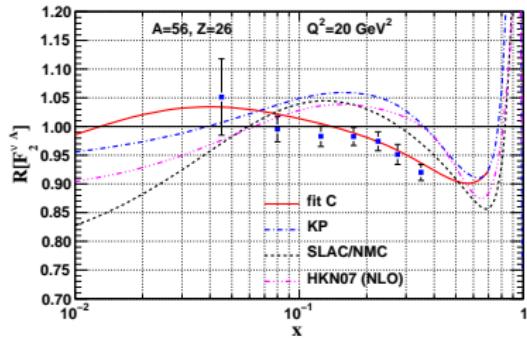
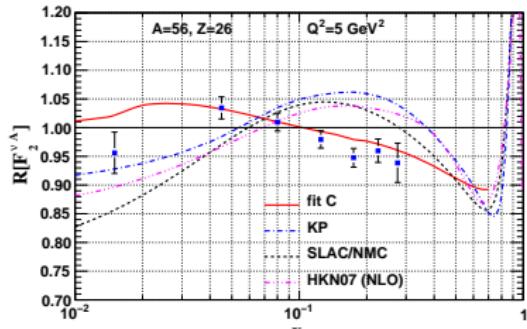
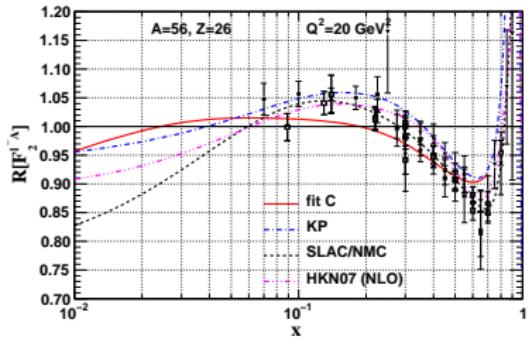
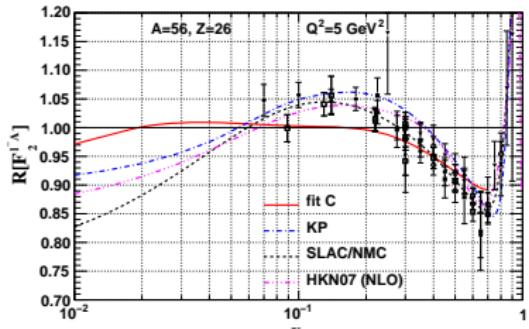
# $R[F_2^{\ell Fe}]$ (LEFT) VS $R[F_2^{\nu Fe}]$ (RIGHT)

global2b ( $w = 1$ )



# $R[F_2^{\ell Fe}]$ (LEFT) VS $R[F_2^{\nu Fe}]$ (RIGHT)

nuanua1 ( $w = \infty$ )



# IS THERE A REASONABLE COMPROMISE FIT?

Weight	Fit name	$\ell$ data	$\chi^2$ (/pt)	$\nu$ data	$\chi^2$ (/pt)	total $\chi^2$ (/pt)
$w = 0$	decut3	708	639 (0.90)	-	-	639 (0.90)
$w = 1/7$	glofac1a	708	645 (0.91)	3134	4710 (1.50)	5355 (1.39)
$w = 1/4$	glofac1c	708	654 (0.92)	3134	4501 (1.43)	5155 (1.34)
$w = 1/2$	glofac1b	708	680 (0.96)	3134	4405 (1.40)	5085 (1.32)
$w = 1$	global2b	708	736 (1.04)	3134	4277 (1.36)	5014 (1.30)
$w = \infty$	nuanua1	-	-	3134	4192 (1.33)	4192 (1.33)

- $w = 0$ : No. Problem:  $R[F_2^{\nu Fe}]$
- $w = 1/7$ : No. Problem:  $R[F_2^{\nu Fe}]$
- $w = 1/4, 1/2$ : No.
  - $Q^2 = 5$ : Undershoots  $R[F_2^{\ell Fe}]$  for  $x < 0.2$ . Overshoots  $R[F_2^{\nu Fe}]$  for  $x \in [0.1, 0.3]$
  - $Q^2 = 20$ :  $R[F_2^{\ell Fe}]$  still ok. Overshoots  $R[F_2^{\nu Fe}]$ .
- $w = 1$ : No. Possibly there is a compromise if more strict  $Q^2$  cut?
  - $Q^2 = 5$ : Undershoots  $R[F_2^{\ell Fe}]$  for  $x < 0.2$ .  $R[F_2^{\nu Fe}]$  ok.
  - $Q^2 = 20$ :  $R[F_2^{\ell Fe}]$  still ok.  $R[F_2^{\nu Fe}]$  ok.
- $w = \infty$ : No. Problem:  $R[F_2^{\ell Fe}]$

# DISCUSSION/INTERMEDIATE CONCLUSION

Discussion based on the comparison of the nuclear correction factors  $R[F_2^{\ell A}]$  and  $R[F_2^{\nu A}]$

- There is definitely a tension between the NuTeV and the charged lepton data
  - There is a clear dependence on the weight.
  - Theory curves for  $R[F_2^{\ell A}]$  and  $R[F_2^{\nu A}]$  are both shifted down with increasing weight of the neutrino data.
- Preliminary conclusion: At the level of the (high) precision there doesn't seem to be a good compromise fit of the combined  $\ell A$ , DY and  $\nu A$  data.
- However one has to be careful:
  - These are precision effects
  - For each weight, the curves have uncertainty bands not considered
  - The figures show the comparison to only few (representative) data

Consider next quantitative criterion based on  $\chi^2$

# TOLERANCE CRITERION

Probability distribution for the  $\chi^2$  function

$$P_N(\chi^2) = \frac{(\chi^2)^{N/2-1} e^{-\chi^2/2}}{2^{N/2} \Gamma(N/2)}$$

Determine  $\xi_{50}^2$  and  $\xi_{90}^2$  (i.e.  $p = 50$ ,  $p = 90$ ):

$$\int_0^{\xi_p^2} d\chi^2 P_N(\chi^2) = p/100$$

Condition for compatibility of two fits:

The 2nd fit ( $\chi_n^2$ ) should be within the 90% C.L. region of the first fit ( $\chi_{n,0}^2$ )

$$\chi_n^2 / \chi_{n,0}^2 < \xi_{90}^2 / \xi_{50}^2 \quad \Leftrightarrow \quad C_{90} \equiv \frac{\Delta \chi^2}{\frac{\chi_{n,0}^2}{\xi_{50}^2} (\xi_{90}^2 - \xi_{50}^2)} < 1$$

see CTEQ'01, PRD65(2001)014012; MSTW'09, EPJC(2009)63,189-285

# TOLERANCE CRITERION $C_{90} < 1$ :

TOTAL  $\chi^2$  FOR A)  $\ell A + \text{DY}$  DATA AND B) NEUTRINO DATA

90% tolerance condition for the **charged lepton**  $\chi^2$  and the **neutrino**  $\chi^2$

- decut3:  $638.9 \pm 45.6$  (best fit to only charged lepton and DY data)
- nuanua1:  $4192 \pm 138$  (best fit to only neutrino data)

Is there a compromise fit compatible to both, decut3 **and** nuanua1?

Weight	Fit name	$\ell$ data	$\chi^2$	$\nu$ data	$\chi^2$	total $\chi^2$ (/pt)
$w = 0$	decut3	708	639	-	nnnn <b>NO</b>	639 (0.90)
$w = 1/7$	glofac1a	708	645 <b>YES</b>	3134	4710 <b>NO</b>	5355 (1.39)
$w = 1/4$	glofac1c	708	654 <b>YES</b>	3134	4501 <b>NO</b>	5155 (1.34)
$w = 1/2$	<b>glofac1b</b>	708	680 <b>YES</b>	3134	4405 <b>NO***</b>	5085 (1.32)
$w = 1$	global2b	708	736 <b>NO</b>	3134	4277 <b>YES</b>	5014 (1.30)
$w = \infty$	nuanua1	-	nnn <b>NO</b>	3134	4192	4192 (1.33)

Observations:

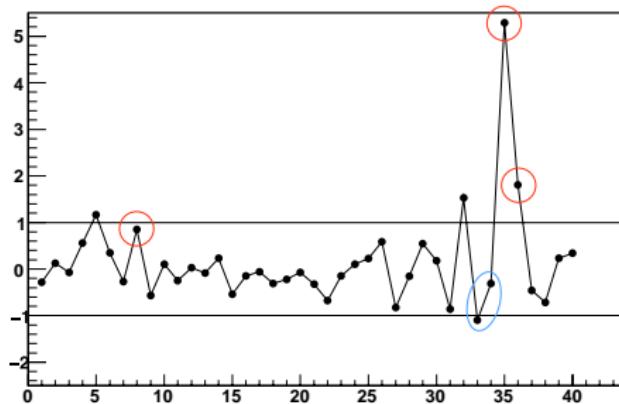
- There is no good compromise fit based on the 90% C.L. criterion.
- Our best candidate is **glofac1b** which is *marginally* compatible:  $4405 - 4192 \simeq 1.5 \times 138$
- Observations in agreement with the previous conclusions based on  $R[F_2^{\ell Fe}]$  and  $R[F_2^{\nu Fe}]$ .

Let's have a look at the tolerance criterion applied to the individual data sets!

# TOLERANCE CRITERION $C_{90} < 1$ :

INDIVIDUAL DATA SETS:  $n = 1, \dots, 32$  VS DECUT3;  $n = 33, \dots, 40$  VS NUANUA1

glofac1a ( $w = 1/7$ )

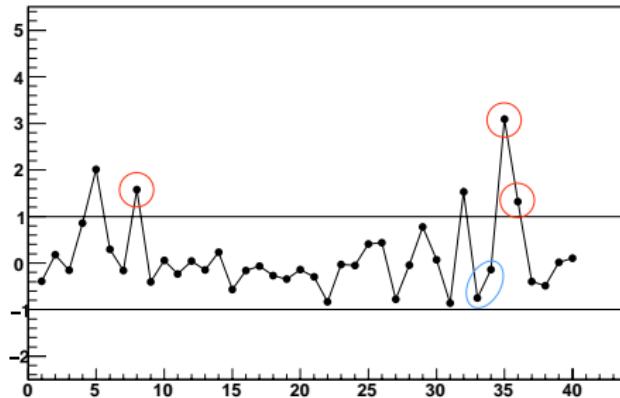


- Y-axis:  $C_{90}$ ; X-axis: Number of the data set ( $n = 1, \dots, 40$ )
- Important data sets:
  - $n = 8$  (red circle):  $Fe/D$  charged lepton data
  - blue ellipse: CHORUS  $\nu Pb$ ,  $\bar{\nu} Pb$  cross section data
  - $n = 35, 36$  (red ellipse): NuTeV  $\nu Fe$ ,  $\bar{\nu} Fe$  cross section data

# TOLERANCE CRITERION $C_{90} < 1$ :

INDIVIDUAL DATA SETS:  $n = 1, \dots, 32$  VS DECUT3;  $n = 33, \dots, 40$  VS NUANUA1

glofac1c ( $w = 1/4$ )

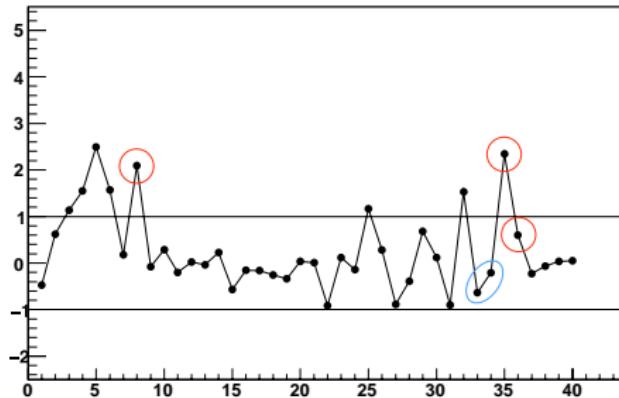


- Y-axis:  $C_{90}$ ; X-axis: Number of the data set ( $n = 1, \dots, 40$ )
- Important data sets:
  - $n = 8$  (red circle):  $Fe/D$  charged lepton data
  - blue ellipse: CHORUS  $\nu Pb, \bar{\nu} Pb$  cross section data
  - $n = 35, 36$  (red ellipse): NuTeV  $\nu Fe, \bar{\nu} Fe$  cross section data

# TOLERANCE CRITERION $C_{90} < 1$ :

INDIVIDUAL DATA SETS:  $n = 1, \dots, 32$  VS DECUT3;  $n = 33, \dots, 40$  VS NUANUA1

glofac1b ( $w = 1/2$ )

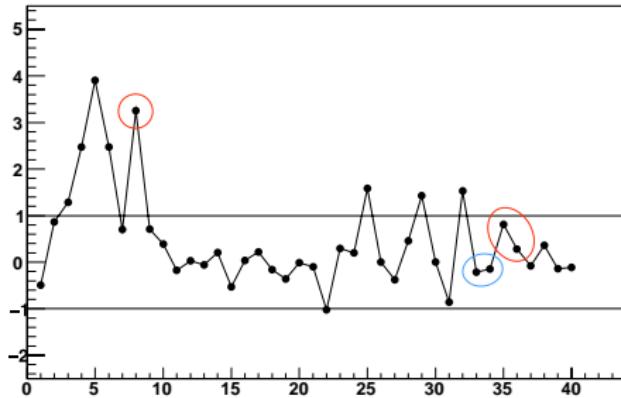


- Y-axis:  $C_{90}$ ; X-axis: Number of the data set ( $n = 1, \dots, 40$ )
- Important data sets:
  - $n = 8$  (red circle):  $Fe/D$  charged lepton data
  - blue ellipse: CHORUS  $\nu Pb$ ,  $\bar{\nu} Pb$  cross section data
  - $n = 35, 36$  (red ellipse): NuTeV  $\nu Fe$ ,  $\bar{\nu} Fe$  cross section data

# TOLERANCE CRITERION $C_{90} < 1$ :

INDIVIDUAL DATA SETS:  $n = 1, \dots, 32$  VS DECUT3;  $n = 33, \dots, 40$  VS NUANUA1

global2b ( $w = 1$ )



- Y-axis:  $C_{90}$ ; X-axis: Number of the data set ( $n = 1, \dots, 40$ )
- Important data sets:
  - $n = 8$  (red circle):  $Fe/D$  charged lepton data
  - blue ellipse: CHORUS  $\nu Pb$ ,  $\bar{\nu} Pb$  cross section data
  - $n = 35, 36$  (red ellipse): NuTeV  $\nu Fe$ ,  $\bar{\nu} Fe$  cross section data

# TOLERANCE CRITERION $C_{90} < 1$ :

## INDIVIDUAL DATA SETS

### Observations:

- $w = 1/7$ :  $C_{90} > 5$  for NuTeV  $\nu Fe$ ;  $C_{90} \simeq 1.8$  for NuTeV  $\bar{\nu} Fe$
- CHORUS data (blue ellipse) always compatible; little dependence on weight  $w$
- increasing weight: NuTeV cross section data improve; charged lepton  $Fe/D$  data get worse
- our best candidate ( $w = 1/2$ )
  - Fe/D ( $n = 8$ ):  $C_{90} \simeq 2$
  - NuTeV  $\nu Fe$  ( $n = 35$ ):  $C_{90} \simeq 2.2$
  - NuTeV  $\bar{\nu} Fe$  ( $n = 36$ ):  $C_{90} < 1$
  - some other data sets  $n = 3, 4, 5, 6, 32$  with  $C_{90} > 1$
- $w = 1$ : Fe/D ( $n = 8$ ):  $C_{90} > 3$
- Confirms and quantifies observations based on  $R$  plots

# CONCLUSIONS

Based on nuclear corrections factors  $R$  and the tolerance criterion  $C_{90} < 1$ :

- There is no good compromise fit to the  $\ell A$  DIS + DY +  $\nu A$  DIS data.
- Most problematic: tension between NuTeV  $\nu Fe$  cross section data and  $Fe/D$  data in charged lepton DIS.
- The NuTeV  $\bar{\nu} Fe$  data are less problematic. They have larger errors.
- The CHORUS  $\nu Pb$  and  $\bar{\nu} Pb$  data are compatible with both, the  $\ell A$ -DIS+DY and the NuTeV  $\nu Fe$  and  $\bar{\nu} Fe$  data, as is well known. They also have larger errors.
- Relaxing the tolerance criterion to  $C_{90} \lesssim 2$  the fit with weight  $w = 1/2$  would be *marginally* acceptable.
- This can also (qualitatively) be verified with the  $R$ -plots.
- A larger  $Q^2$ -cut, say  $Q^2 > 5 \text{ GeV}^2$ , could also help to reduce the tension. (In particular, this would remove some of the rather precise NuTeV cross section data at small  $x$ .)